# Astrometric Measurements of Seven Double Stars, September 2011 Report 

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#### Abstract

From my residence in Paso Robles, California, measurements of the separation and position angle of seven double stars were made. Listed in chronological order, the double stars were Zeta Ursae Majoris, Zeta Lyrae, Epsilon Delphini, SAO 105104 in Sagitta, STF 2840 in Cepheus, 61 Cygni, and 17 Cygni. The two goals of this project were to measure the position angle and separation of the aforementioned double stars, and to learn the necessary techniques to conduct this research.


## Methodology

My observations were made from my home in Paso Robles, California (located at approximately $35^{\circ} 37^{\prime} 36^{\prime \prime} \mathrm{N}$ and $\left.120^{\circ} 41^{\prime} 24^{\prime \prime} \mathrm{W}\right)$ using a Celestron model CPC 1100 telescope (Figure 1). The telescope is computerized, motorized, and was fitted with a Celestron Micro Guide 12.5 mm astrometric eyepiece. The telescope is of Schmidt-Cassegrain design, with aperture of 11 inches on an alt-azimuth mount. The manufacturer reports a focal length of $2,800 \mathrm{~mm}$.

The Micro Guide eyepiece was oriented with the celestial coordinate system using the primary star of the double star under study. The primary star was positioned on the mark 30 , the drive was disabled, and the star was permitted to drift to the outer circle. The scale was rotated until the star lay on the 270 degree mark. The accuracy of this setting was verified by positioning the primary star on the 90 degree mark of the outer circular scale, and allowing the star to drift to the 270 degree mark.

Following the orientation, drift times were measured by placing the primary star on the 0 mark of the linear scale, and measuring the drift time from the 0 to the 60 mark using a stop watch precise to $\pm 0.01$ seconds. Measurements were made, and the average drift time was calculated. That average was used to calculate the scale constant Z , using the formula ${ }^{14}$ :


Figure 1: The author with his Celestron telescope

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$$
Z=\frac{15.0411 T_{\text {ave }} \cos \delta}{D}
$$

where $T_{\text {ave }}$ is the average time, $\delta$ is the declination angle, and $D$ is the number of reticle divisions.

Separation measurements were made by placing the pair of stars on the linear scale at the zero mark, and then counting the number of scale divisions between the stars. Because the scale has 60 divisions, it was only possible to estimate to the nearest $1 / 4$ division. After each measurement, the double star was repositioned to the next major division. Measurements were made, and an average and standard deviation were calculated.

The position angle measurements were made by aligning both stars on the linear scale with the primary star at the 30 division and pointing to the 60 mark, disabling the tracking feature, and then allowing the stars to drift to the circular scales. The crossing of the primary star at the outer scale was approximated to the nearest degree as the scale has divisions of $5^{\circ}$. Following each measurement, the tracking feature was enabled and the process was repeated.

## Zeta Ursae Majoris - Introduction

This double star is located in the constellation of Ursa Major (the Great Bear), and is known by its traditional name of Mizar with alternate spellings of Mirzar, Mizat, and Mirza. This double star has been known since ancient times, and was the first double star. On clear nights, the double star can be seen without the use of instruments. It was studied by Benedetto Castelli in 1617, and has been studied frequently since that time. Both stars are yellow with magnitudes of 2.0 and $4.0^{22}$. The colors reported for this pair vary considerably having been reported as both white, white and emerald, both green, blue, or yellow. For the AB components, Right Ascension is $13^{\mathrm{h}} 23^{\mathrm{m}} 56^{\mathrm{s}}$ and the Declination is $+544^{\circ} 55^{\prime} 31^{\prime \prime}$. What was once thought to be a double star is actually a complex of six stars which are all gravitationally bound ${ }^{20}$.

The catalog identifiers for this double star include 79 Ursae Majoris, ADS 8891AB, BD+55 1598A, BGC 18133, FK5 497, HD 116656, HIP 65378, HR 5054, SAO 28737, STF 1744, and WDS 13239+5456. Its precise coordinates are $132355.42+545531.5 .{ }^{22}$

## Zeta Ursae Majoris - Observations

The measurements were made on 11 May 2011 (Bessell date 2011.359) beginning at 9:50pm and end-

Table 1: Literature Search; Separation (arc seconds); Position angle (degrees) for Zeta Ursae Majoris

| Reference name | Sep | PA |
| :--- | :--- | :--- |
| William Herschel Catalog 1779 data $^{19}$ | 14.3 | 153 |
| Washington Double Star Cat. $^{35} 1993$ data | 14.5 | 152 |
| Eagle Creek Observatory ${ }^{24}$ | 14.4 | 151 |
| Daley, $2009^{37}$ | 14.3 | 152.5 |
| Measurements by the author 2011 | 14.7 | 152.3 |

ing at 11:50pm Pacific Daylight Time. The night was clear and calm, and there was a $1 / 2$ moon. The temperature ranged from 60 to $50^{\circ} \mathrm{F}$. There was a breeze of $5-10 \mathrm{mph}$ which affected the telescope and several measurements were repeated.

The linear scale of the Micro Guide eyepiece was oriented with the celestial coordinate system using the primary star. Once the orientation was completed, 12 drift time measurements were made, with an average value of 47.19 seconds, a standard deviation of 0.63 seconds, and a standard error of the mean of 0.18 seconds. The result was a scale constant of 6.8 arc seconds per division.

The primary star was placed on the linear scale, and 12 separation measurements were taken. The average value was 2.2 divisions with a standard deviation of 0.22 divisions, and a standard error of the mean of 0.06 divisions. When adjusted for significant figures, the calculated separation was 14.7 arc seconds.

The position angle measurements were made using the aforementioned methodology, and 18 position angle measurements were taken with an average value of $152.3^{\circ}$, a standard deviation of $1.5^{\circ}$, and a standard error of the mean of $0.27^{\circ}$.

The separation value of 14.7 arc seconds and position angle value of $152.3^{\circ}$ compared well with the separation value of 14.5 arc seconds and the position angle value of $152^{\circ}$ published in the Washington Double Star catalog. ${ }^{35}$ See Table 1.

## Zeta Lyrae - Introduction

Located in the constellation of Lyra (the Harp), the double star Zeta Lyrae has been known since ancient times. Its right ascension is $18^{\mathrm{h}} 44^{\mathrm{m}} 46^{\mathrm{s}}$ and its declination is $+377^{\circ} 36^{\prime} 18^{\prime \prime}$. The yellow primary and blue-white secondary stars have magnitudes of 4.4 and 5.7 respectively. Modern observations have

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shown that Zeta Lyrae is a spectroscopic binary (Wikipedia).

The catalog identifiers for this pair include 6 Lyrae, $\mathrm{BD}+37$ 3222, BU 968, CSV 101763, GSC 03118+02080, HD 173648, HIP 91971, HR 7056, PPM 81740, SAO 67321, STF 38, UBV 15954, and WDS $18448+3736$. Its precise coordinates are $184446.34+373618.2 .{ }^{22}$

## Zeta Lyrae - Observations

The measurements were made on 11 July 2011 (Bessell date 2011.526) beginning at $10: 15 \mathrm{pm}$ and ending at $11: 50 \mathrm{pm}$ Pacific Daylight Time. The night was clear, and the moon was gibbous. The temperature ranged from 60 to $50^{\circ} \mathrm{F}$. The wind was gentle at $1-5 \mathrm{mph}$.

The linear scale of the Micro Guide eyepiece was oriented with the celestial coordinate system using the primary star. Once the orientation was completed, 12 drift time measurements were made, with an average value of 35.94 seconds, a standard deviation of 0.29 seconds, and a standard error of the mean of 0.08 seconds. The result was a scale constant of 7.13 are seconds per division.

The primary star was placed on the linear scale, and 12 separation measurements were taken. The average value was 6.17 divisions with a standard deviation of 0.25 divisions, and a standard error of the mean of 0.07 divisions. When adjusted for significant figures, the calculated separation was 44.0 arc seconds.

The position angle measurements were made using the aforementioned methodology, and 18 position angle measurements were taken with an average value of $149.9^{\circ}$, a standard deviation of $2.18^{\circ}$, and a standard error of the mean of $0.4^{\circ}$.

The separation value of 44.0 arc seconds and position angle value of $149.9^{\circ}$ compared well with the separation value of 43.7 arc seconds and the position angle value of $150^{\circ}$ published in the Washington Double Star catalog ${ }^{22}$. See Table 2.

## Epsilon Delphini - Introduction

Located in the constellation of Delphinus (the Dolphin), the double star Epsilon Delphini consists of a pair of yellow-white stars of magnitudes 7.1 and 7.4. For the AB components, the right ascension is $20^{\mathrm{h}} 31^{\mathrm{m}} 12^{\mathrm{s}}$ and its declination is $+11^{\circ} 15^{\prime} 34^{\prime \prime}{ }^{22}$. It does not have a traditional name.

The catalog identifiers for this pair include ADS 13946BC, AG+11 2505, BD+10 4307B, CSI+10 4307 1, GC 28544, GCRV 12819, GEN+1.00195482, HD

Table 2 Literature Review: Separation (arc seconds) and Position angle (degrees) for Zeta Lyrae

| Reference name | Sep | PA |
| :--- | :--- | :--- |
| Washington Double Star Cat. ${ }^{35} 1835$ data | 43.7 | 150 |
| Bright Star Catalog $^{17}$ | 43.7 | - |
| Burton, 2006 ${ }^{10}$ | 43.6 | 150 |
| Perez, 2005 ${ }^{27}$ | 43.7 | 150 |
| Arnold, 2008 |  |  |
| Bell, 2011 | 43.9 | 149.8 |
| Schlimmer ${ }^{33} 2009$ data | 44 | 150 |
| Washington Double Star Cat. ${ }^{22} 2010$ data | 43.7 | 150 |
| Measurements by the author 2011 | 44.0 | 149.8 |

195482, HIP 101233, IDS 20264+1055, PPM 138507, SAO 106195, SKY 38830, STF 2690, UBV M24917, and WDS J20312+1116BC. Its precise coordinates are 203111.94+111533.7. ${ }^{22}$

## Epsilon Delphini - Observations

The measurements were made on 1 August 2011 (Bessell date 2011.584) beginning at 9:30pm and ending at 11:00pm Pacific Daylight Time. The night was clear, with no moon. The temperature range was from 65 to $55^{\circ} \mathrm{F}$. There was a $0-5 \mathrm{mph}$ breeze, and the humidity was $25 \%$.

The linear scale of the Micro Guide eyepiece was oriented with the celestial coordinate system using the primary star. Once the orientation was completed, 12 drift time measurements were made, with an average value of 28.91 seconds, a standard deviation of 0.28 seconds, and a standard error of the mean of 0.08 seconds. The result was a scale constant of 7.11 arc seconds per division.

The primary star was placed on the linear scale, and 12 separation measurements were taken. The average value was 2.5 divisions with a standard deviation of 0.0 divisions, and a standard error of the mean of 0.0 divisions. When adjusted for significant figures, the calculated separation was 17.7 arc seconds.

The position angle measurements were made using the aforementioned methodology, and 18 position angle measurements were taken with an average value of $256^{\circ}$, a standard deviation of $1.5^{\circ}$, and a standard error of the mean of $0.28^{\circ}$.

The separation value of 17.7 arc seconds and the position angle of $256^{\circ}$ compared well with the values

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Table 3: Literature Review, Separation (arc seconds) and Position angle (degrees) for Epsilon Delphini

| Reference name | Sep | PA |
| :--- | :--- | :--- |
| Washington Double Star Cat. ${ }^{22} 1777$ data | 15.0 | 281 |
| The Hipparcos Catalog $^{28} 1997$ data | 16.7 | 253 |
| Eagle Creek Observatory ${ }^{24}$ | 16.7 | 256 |
| Arnold, $2006^{5}$ | 17.8 | 254.8 |
| Schlimmer, $2007^{30}$ | 17.3 | 254.1 |
| Schlimmer, $2009^{32}$ | 17.7 | 254.8 |
| Washington Double Star Cat. ${ }^{22} 2009$ data | 17.3 | 255 |
| Schlimmer 2010 ${ }^{33}$ | 17.3 | 255.1 |
| The Tycho Catalog ${ }^{15}$ | 17.4 | 254.8 |
| Measurements by the author 2011 | 17.7 | 255.9 |

of 17.3 and $255^{\circ}$ as given in the Washington Double Star catalog ${ }^{22}$. See Table 3.

## SAO 105104 in Sagitta - Introduction

Located in the constellation of Sagitta (the Arrow) is this double star of which the primary star is yelloworange and the secondary star is white. With magnitudes of 6.4 and 9.5 , the secondary star approached the limit of the CPC 1100 telescope. Its Right Ascension is $19^{\mathrm{h}} 39^{\mathrm{m}} 25^{\mathrm{s}}$ and its Declination is $+16^{\circ} 34^{\prime} 16^{\prime \prime 22}$. This double star has no traditional name.

The catalog identifiers includeAG+16 2018, BC+16 3936, GSC 01602-01582, HD 185622, JIP 96688, HP 7475, IRAS 19371+1627, PPM 136711, SAO 105104, TYC 1602-1582-1, V 0340, and WDS 1934+1634A. Its precise coordinates are $193925.33+163416.0 .{ }^{22}$

## SAO 105104 in Sagitta - Observations

These measurements were made on 3 August 2011 (Bessell date 2011.589) beginning at 9:30pm and ending at 11:00pm Pacific Daylight Time. The night was clear, with a $1 / 4$ moon in the southwest. The temperature range was from 65 to $55^{\circ} \mathrm{F}$. There was a 0 5 mph breeze. The humidity was $30 \%$.

The linear scale of the Micro Guide eyepiece was oriented with the celestial coordinate system using the primary star. Once the orientation was completed, 12 drift time measurements were made, with an average value of 30.36 seconds, a standard deviation of 0.15 seconds, and a standard error of the mean of 0.04 seconds. The result was a scale constant of 7.3

Table 4: Literature Review, Separation (arc seconds); Position angle (in ${ }^{\circ}$ ) for SAO 105104 in Sagitta

| Reference name | Sep | PA |
| :--- | :--- | :--- |
| W. Herschel (McEvoy 2011) ${ }^{19} 1796$ data | 28.2 | 302 |
| Hipparcos Catalog $^{28}$ | 30.0 | 301.2 |
| SKY2000 Master Catalog $^{25}$ | 28.2 | 302 |
| C C D M Catalog |  |  |
| Arnold 2010 | 28.2 | 302 |
| Washington Double Star Cat. ${ }^{22} 2010$ data | 28.6 | 301 |
| Measurements by the author 2011 | 29.0 | 301 |

arc seconds per division.
The primary star was placed on the linear scale, and 12 separation measurements were taken. The average value was 4.0 divisions with a standard deviation of 0.0 divisions, and a standard error of the mean of 0.0 divisions. When adjusted for significant figures, the calculated separation was 29.0 arc seconds.

The position angle measurements were made using the aforementioned methodology, and 18 position angle measurements were taken with an average value of $301^{\circ}$, a standard deviation of $1.7^{\circ}$, and a standard error of the mean of $0.31^{\circ}$.

The separation value of 29.0 arc seconds and position angle value of $301^{\circ}$ compared well with the separation value of 28.6 arc seconds and the position angle value of $301^{\circ}$ published in the Washington Double Star Catalog ${ }^{22}$. See Table 4.

## STF 2840 in Cepheus - Introduction

Located in the constellation of Cepheus, this pair of yellow stars has magnitudes of 5.6 and 6.4. STF 2840 in Cepheus has been well studied, but little has been written about this pair. The WDS gives its Right Ascension is $21^{\mathrm{h}} 52^{\mathrm{m}} 01^{\mathrm{s}}$ and its Declination is $+55^{\circ} 47^{\prime} 48^{\prime \prime 2}$, however, the Cambridge Double Star Atlas lists an R.A. of $21^{\mathrm{h}} 52^{\mathrm{m}} 19 \mathrm{~s}$ and a Dec. of $+55^{\circ} 50^{\prime} 10^{\prime \prime}$, and the data from the Hipparcos Catalog ${ }^{28}$ is an R.A. of. $21^{\mathrm{h}} 52^{\mathrm{m}} 00^{\mathrm{s}}$ and a Dec. of $+55^{\circ} 47^{\prime} 31^{\prime \prime}$.

The catalog identifiers include ADS 15045, AG+55 1503, BD+55 2638, HD 208063, HIP 107929, IDS 21486+5519B, PPM 39938, SAO 33817, SKY 41670, TYC 3972-2737-1, UBV 187420, and WDS $21520+5548$ B. Its precise coordinates are $215201.02+554748.3^{22}$

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Table 5: Literature Review: Separation (arc seconds); Position angle (degrees) for STF 2840 in Cepheus

| Reference name | Sep | PA |
| :--- | :--- | :--- |
| Washington Double Star Cat. ${ }^{22} 1782$ data | 21.2 | 192 |
| Hipparcos Catalog $^{28} 1993$ data | 17.9 | 197 |
| Eagle Creek Observatory ${ }^{24}$ | 18.3 | 196 |
| Arnold $2010^{8}$ | 17.8 | 196.8 |
| Washington Double Star Cat. ${ }^{22} 2010$ data | 17.8 | 197 |
| Measurements by the author 2011 | 18.1 | 196 |

## STF 2840 in Cepheus - Observations

The measurements were made on 9 August 2011 (Bessell date 2011.605) beginning at $10: 20 \mathrm{pm}$ and ending at 11:30 pm Pacific Daylight Time. The night was clear, calm, with a gentle breeze of $1-5 \mathrm{mph}$. The moon was $2 / 3$ full in the southwest. The temperature ranged from 60 to $50^{\circ} \mathrm{F}$. The humidity was $65 \%$, and seeing was $3-4$.

The linear scale of the Micro Guide eyepiece was oriented with the celestial coordinate system using the primary star. Once the orientation was completed, 12 drift time measurements were made, with an average value of 51.55 seconds, a standard deviation of 0.25 seconds, and a standard error of the mean of 0.08 seconds. The result was a scale constant of 7.3 arc seconds per division.

The primary star was placed on the linear scale, and 12 separation measurements were taken. The average value was 2.5 divisions with a standard deviation of 0.0 divisions, and a standard error of the mean of 0.0 divisions. When adjusted for significant figures, the calculated separation was 18.1 arc seconds.

The position angle measurements were made using the aforementioned methodology, and 18 position angle measurements were taken with an average value of $196^{\circ}$, a standard deviation of $1.3^{\circ}$, and a standard error of the mean of $0.24^{\circ}$.

The separation value of 18.1 arc seconds and position angle value of $196^{\circ}$ compared well with the separation value of 17.8 arc seconds and the position angle value of $197^{\circ}$ published in the Washington Double Star catalog ${ }^{22}$. See Table 5.

## 61 Cygni - Introduction

Located in the constellation of Cygnus (the Swan) is 61 Cygni, a famous pair of orange stars. First studied Piazzi in 1792, the pair has a large proper motion

Table 6: Literature Review, Separation (arc seconds) and Position angle (degrees) for 61 Cygni

| Reference name | Sep | PA |
| :---: | :---: | :---: |
| W. Herschel (MacEvoy 2011) ${ }^{19} 1753$ data | 35 | 19.6 |
| Bright Star Cat. ${ }^{17}$ | 28.7 | - |
| Washington Double Star Cat. ${ }^{35}$ 1993 data | 30.4 | 149 |
| The Hipparchos Catalog ${ }^{28} 1993$ data ${ }^{22}$ | 31.6 | 151.2 |
| Arnold $2006{ }^{4} 2004$ data | 31.1 | 150.9 |
| Muller $2007^{23} 2005$ data | 31.6 | 152.6 |
| Perez 2006 ${ }^{27}$ | 36 | 140 |
| Schlimmer $2009^{31} 2008$ data | 31.0 | 151.3 |
| Heijen $2008{ }^{16}$ | 31.1 | 151 |
| Anton $2011{ }^{2}$ Oct 9 | 31.2 | 151.3 |
| Anton $2011{ }^{2}$ Oct 12 | 31.35 | 151.6 |
| Washington Double Star Cat. ${ }^{22}$ 2010 data | 31.4 | 152 |
| Measurements by the author 2011 | 31.9 | 151 |

of about 5 arc seconds per year, and is sometimes called the "Flying Star"3. In 1838 Bessell measured the parallax and distance from the Earth for this pair, which was the first double star to be so measured ${ }^{38}$. The orange pair is distinct, but the surroundings lack any prominent stars. The pair has magnitudes of 5.2 and 6.1. Its Right Ascension is $21^{\mathrm{h}} 06^{\mathrm{m}} 54^{\mathrm{s}}$ and its Declination is $+38^{\circ} 44^{\prime} 58^{\prime \prime}{ }^{22}$.

The catalog identifiers include BD+38 4343, FK5 793, HD 201091, HIP 104214, HR 8085, GC 29509, GJ 820, PPM 86045, SAO 70919, STF 2758, UBV 18287, and WDS 21069+3845. Its precise coordinates are $210653.94+384457.8 .{ }^{22}$

## 61 Cygni - Observations

The measurements were made on 14 August 2011 (Bessell date 2011.619) beginning at $8: 45 \mathrm{pm}$ and ending at 9:30pm Pacific Daylight Time. The night was clear and calm with no moon. The temperature ranged from 75 to $65^{\circ} \mathrm{F}$. The humidity was $25 \%$, and seeing at $4-5$. The wind at $5-10 \mathrm{mph}$ affected the measurements, and many were repeated.

The linear scale of the Micro Guide eyepiece was oriented with the celestial coordinate system using the primary star. Once the orientation was completed, 12 drift time measurements were made, with an average value of 36.17 seconds, a standard devia-

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Table 7: Literature Review, Separation (arc seconds) and Position angle (degrees) for 17 Cygni

| Reference name | Sep | PA |
| :--- | :--- | :--- |
| W. Herschel (MacEvoy 2011) ${ }^{19} 1822$ data | 25.5 | 73 |
| Starland Catalog (Olcott 1909$)^{26}$ | 26.2 | 69 |
| Washington Double Star Cat. ${ }^{35} 1993$ data | 26.3 | 69 |
| Hipparcos Catalog ${ }^{28}$ | 26.2 | 67.5 |
| Eagle Creek Observatory ${ }^{24}$ | 26 | 70 |
| Astrogeek (Burton 2011) ${ }^{10} 2006$ data | 26.3 | 67 |
| Arnold 2010 | 26.2 | 69 |
| Washington Double Star Cat. ${ }^{22} 2009$ data | 26.2 | 69 |
| Measurements by the author 2011 | 25.5 | 69 |

tion of 0.37 seconds, and a standard error of the mean of 0.09 seconds. The result was a scale constant of 7.07 arc seconds per division.

The primary star was placed on the linear scale, and 12 separation measurements were taken. The average value was 4.5 divisions with a standard deviation of 0.1 divisions, and a standard error of the mean of 0.03 divisions. When adjusted for significant figures, the calculated separation was 31.9 arc seconds.

The position angle measurements were made using the aforementioned methodology, and 18 position angle measurements were taken with an average value of $150.8^{\circ}$, a standard deviation of $1.3^{\circ}$, and a standard error of the mean of $0.31^{\circ}$.

The separation value of 31.9 arc seconds and the position angle value of 151 degrees compared favorably with the values of 31.4 arc seconds and 152 degrees as published in the Washington Double Star Catalog (Mason+2011) ${ }^{22}$. See Table 6.

## 17 Cygni - Introduction

Located in the constellation of Cygnus (the Swan), 17 Cygni is a yellow pair of stars with magnitudes of 5.1 and 9.3. Its right ascension is $19^{\mathrm{h}} 46^{\mathrm{m}} 26^{\mathrm{s}}$ and its declination is $+33^{\circ} 43 \prime 39^{\prime \prime} .^{22}$

The catalog identifiers include ADS 12913A, BD+33 3587, CSI+33 3587 1, GC 27369, HD 187013, HIP 97295, HR 7534, PLX 4654, PPM 83516, SAO 68827, STF 2580AB, TYC 2660-4227-1, and WDS $19464+3344$. Its precise coordinates are $194625.60+334339.3 .{ }^{22}$

## 17 Cygni - Observations

The measurements were made on 15 August 2011 (Bessell date 2011.622) beginning at $8: 30 \mathrm{pm}$ and ending at 10:30 pm Pacific Daylight Time. The night was clear with moonrise at $9: 30 \mathrm{pm}$. The temperature ranged from 75 to $65^{\circ} \mathrm{F}$. The wind was $1-5 \mathrm{mph}$, humidity $30 \%$, and seeing $3-4$.

The linear scale of the Micro Guide eyepiece was oriented with the celestial coordinate system using the primary star. Once the orientation was completed, 12 drift time measurements were made, with an average value of 33.84 seconds, a standard deviation of 0.2 seconds, and a standard error of the mean of 0.06 seconds. The result was a scale constant of 7.1 arc seconds per division.

The primary star was placed on the linear scale, and 12 separation measurements were taken. The average value was 3.6 divisions with a standard deviation of 0.10 divisions, and a standard error of the mean of 0.03 divisions. When adjusted for significant figures, the calculated separation was 25.5 arc seconds.

The position angle measurements were made using the aforementioned methodology, and 18 position angle measurements were taken with an average value of $68.9^{\circ}$, a standard deviation of $1.35^{\circ}$, and a standard error of the mean of $0.25^{\circ}$.

The separation value of 25.5 arc seconds and the position angle value of 69 degrees compared well with the values of 26 arc seconds and 69 degrees from the Washington Double Star Catalog ${ }^{22}$. See Table 7.

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